

# ARIEL

Enabling planetary science across light-years



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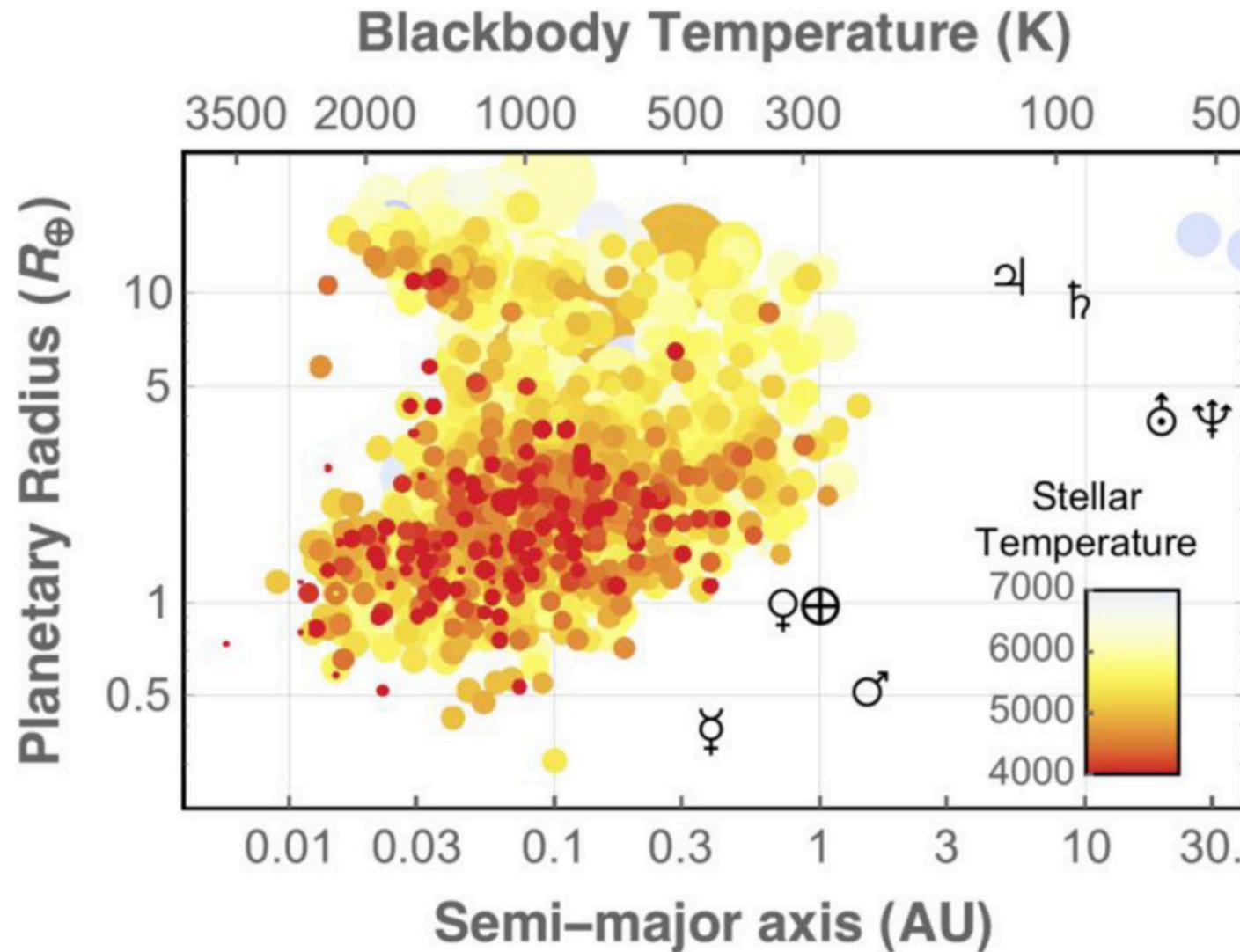
Institut de Ciències de l'Espai (ICE, CSIC)

1st meeting of ExoNet, Granada, 22/2/2019



# Exoplanets today: huge diversity

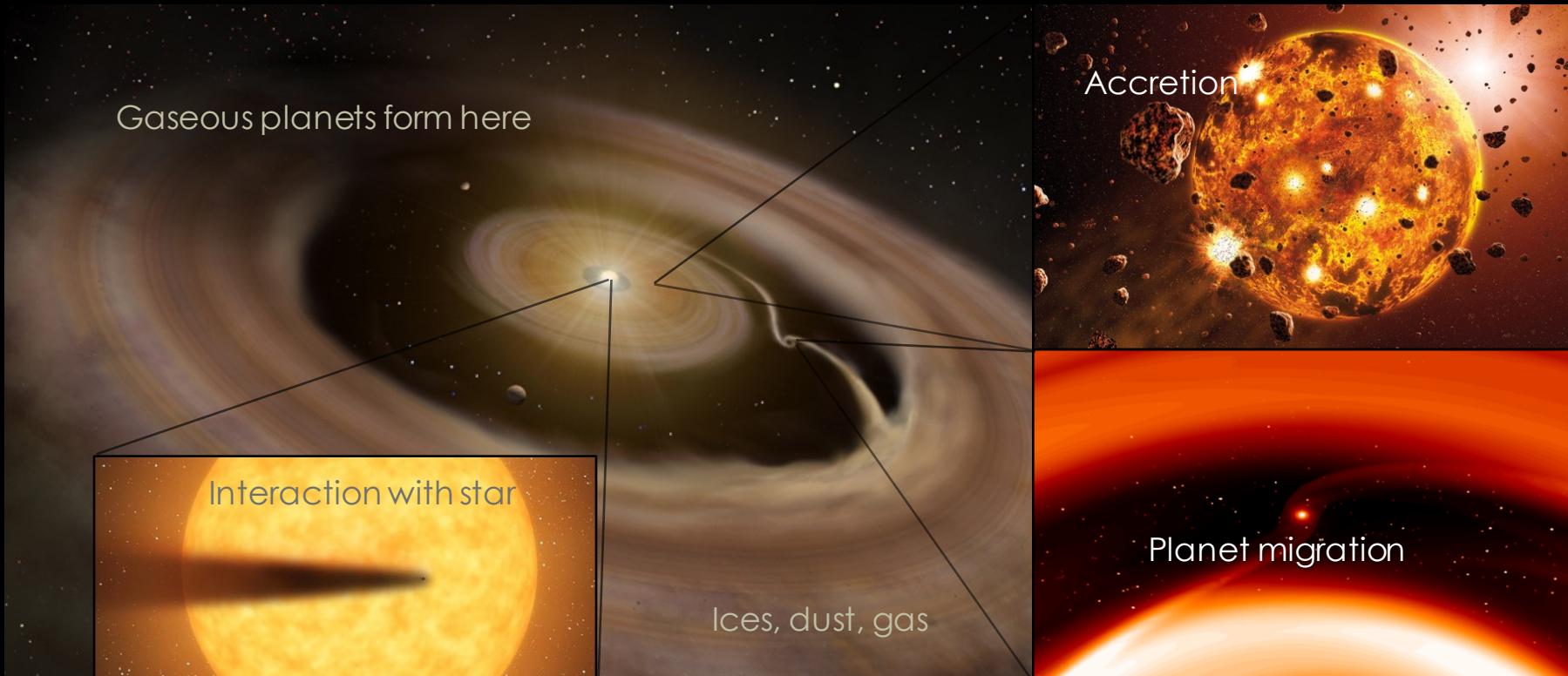
3800+ PLANETS, 2700 PLANETARY SYSTEMS KNOWN IN THE GALAXY



# Huge diversity: why?



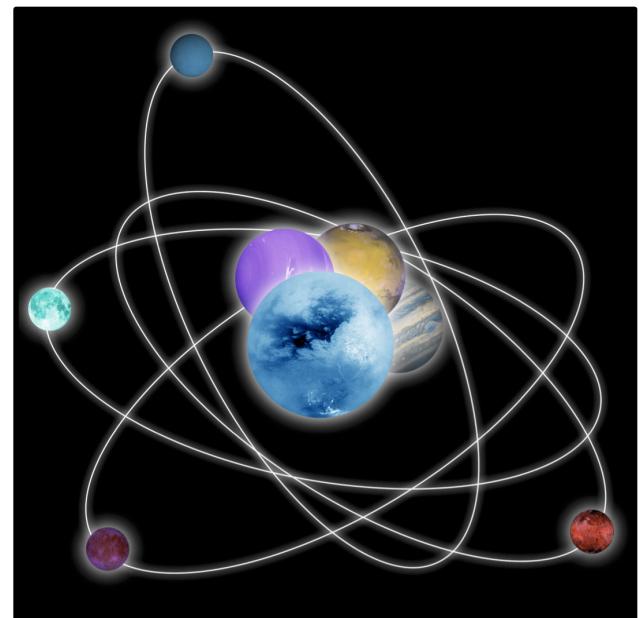
Formation & evolution processes?  
Migration? Interaction with star?



# *Key exoplanet questions*



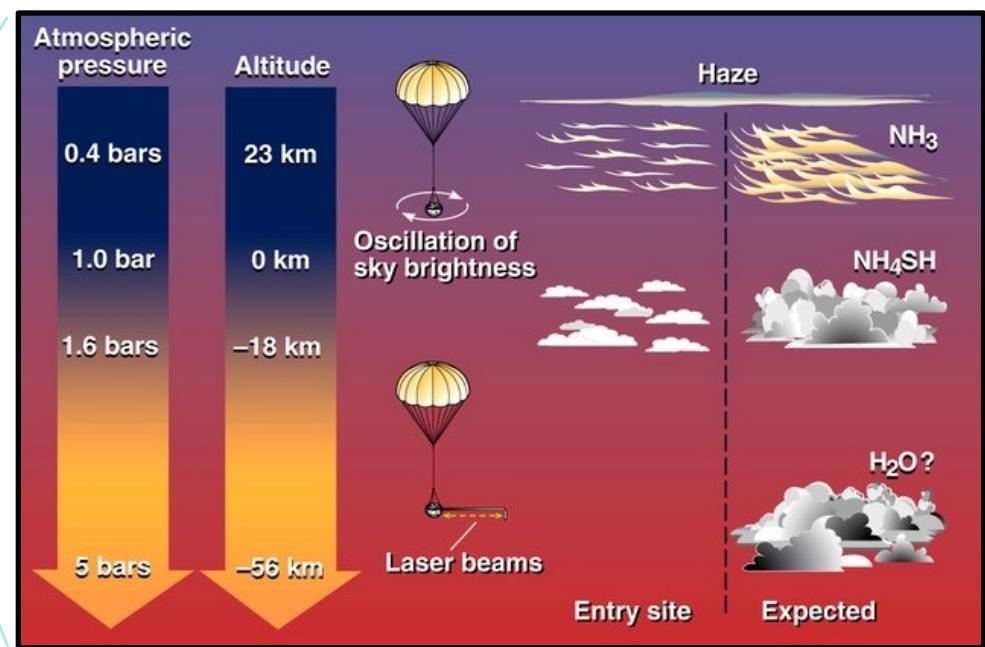
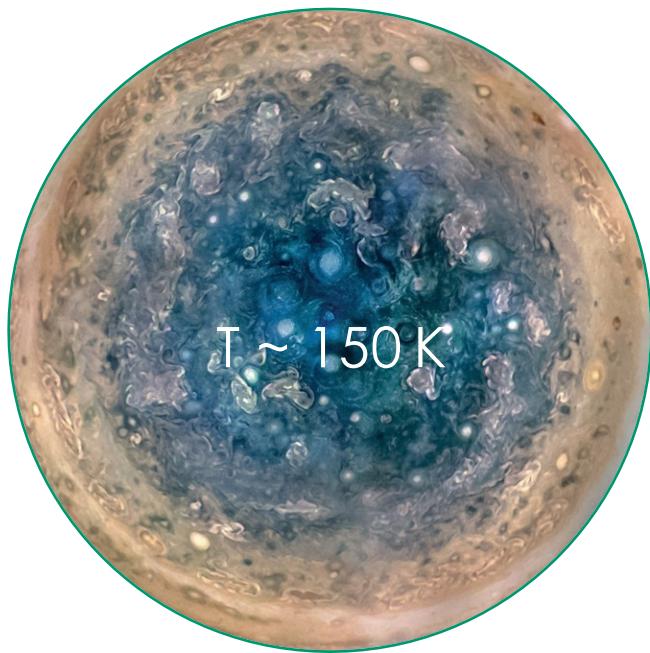
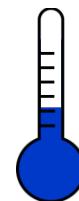
- How diverse are exoplanets chemically?
- Does chemical diversity correlate with other parameters?
  - How do planets form?
  - How do planets evolve?





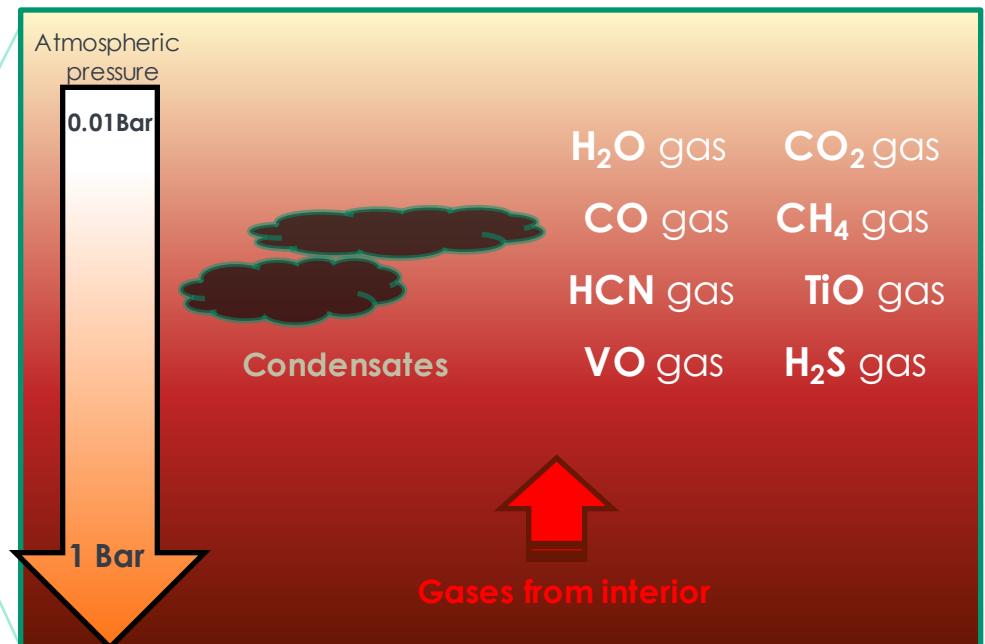
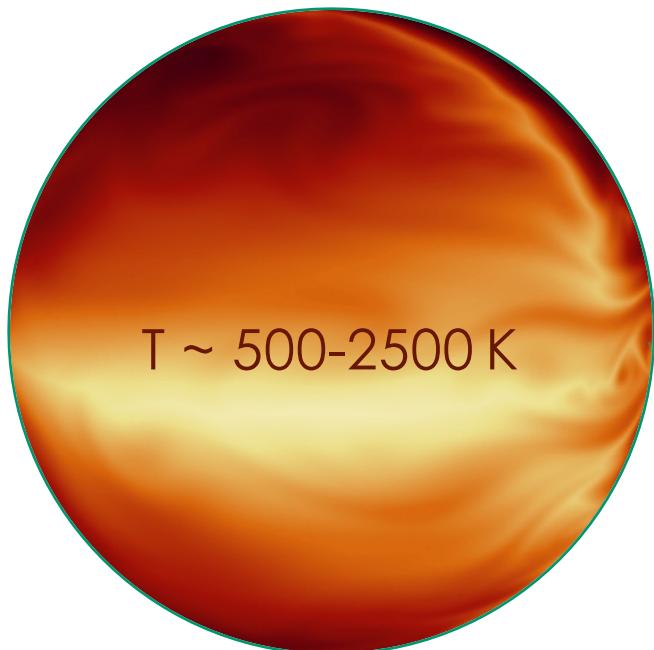
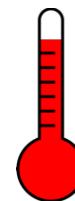
# The Sun's planets are cold

SOME KEY O, C, N, S MOLECULES ARE **NOT** IN GAS FORM



# Warm/hot exoplanets

O, C, N, S (Ti, VO, Si) MOLECULES ARE IN GAS FORM





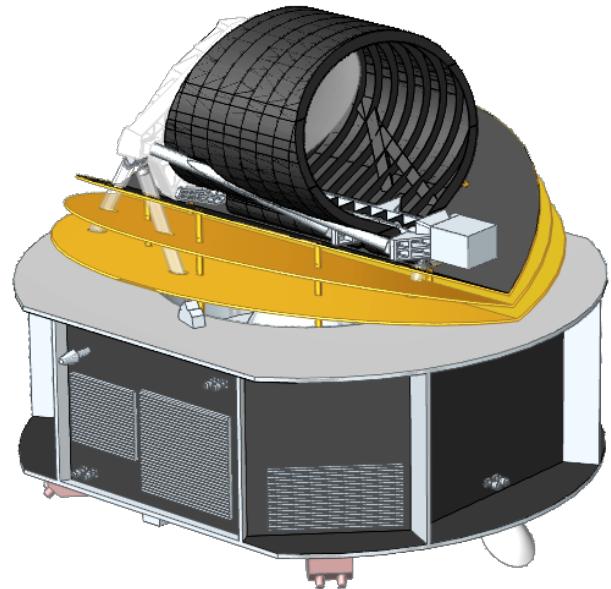
# ***Issues with current data***

- Low SNR & R observations
- Data are sparse, not enough wavelength coverage
- Broad wavelength coverage is not simultaneous
- Absolute calibration at the level of  $10^{-4}$  is not guaranteed!
- Instrument systematics are difficult to disentangle from the signal
- Stellar activity is the largest source of astrophysical noise
- We need observations on a population of objects to draw conclusions

# *ARIEL – ESA M4 mission*

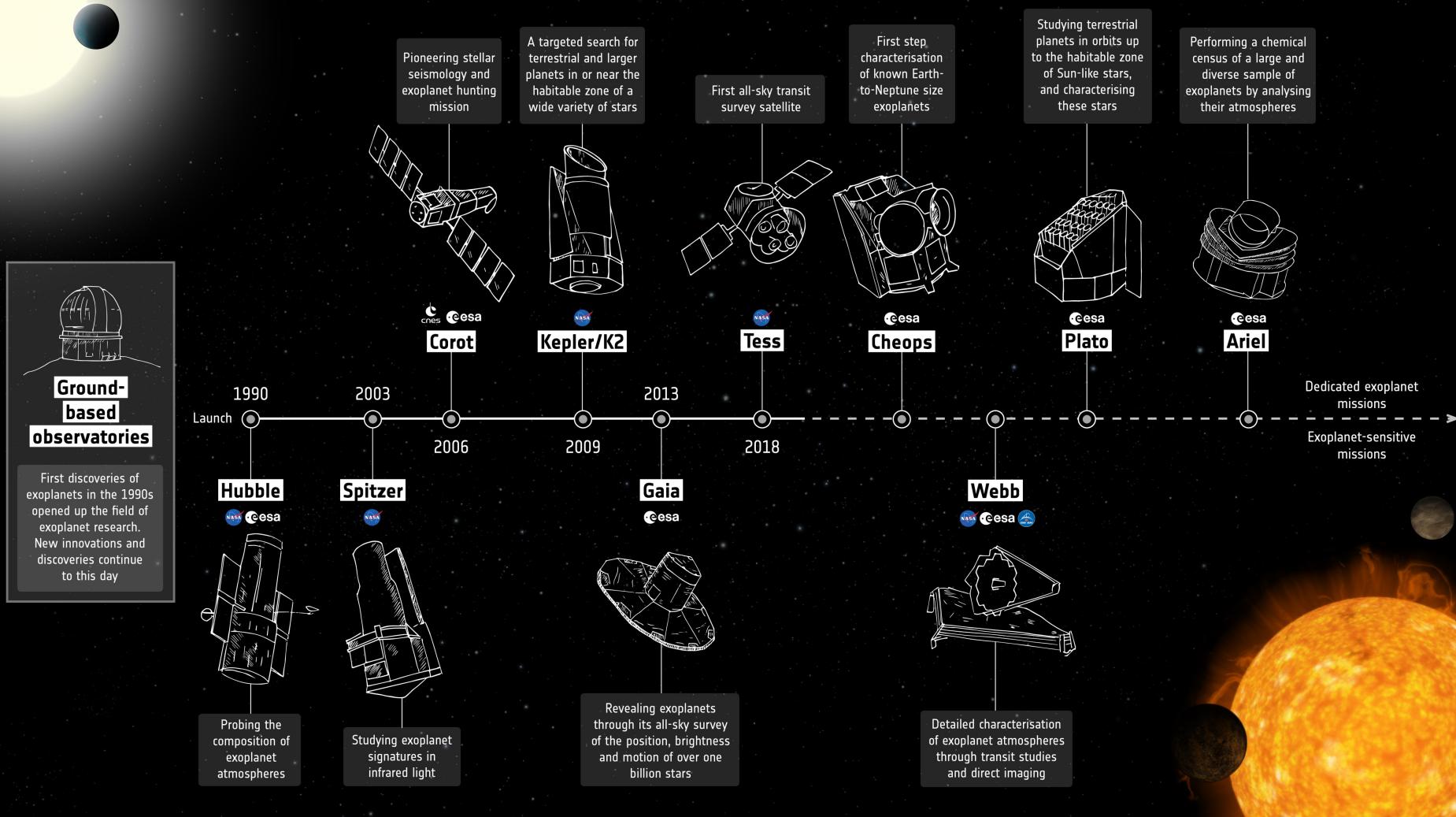


- 1-m telescope, spectroscopy from VIS to IR
- Satellite in orbit around L2
- ~1000 exoplanets observed (rocky + gaseous)
- Launch in 2028 – 3.5-yr mission
- Simultaneous coverage 0.5-7.8 micron
- Payload consortium: 15 ESA countries + NASA under study



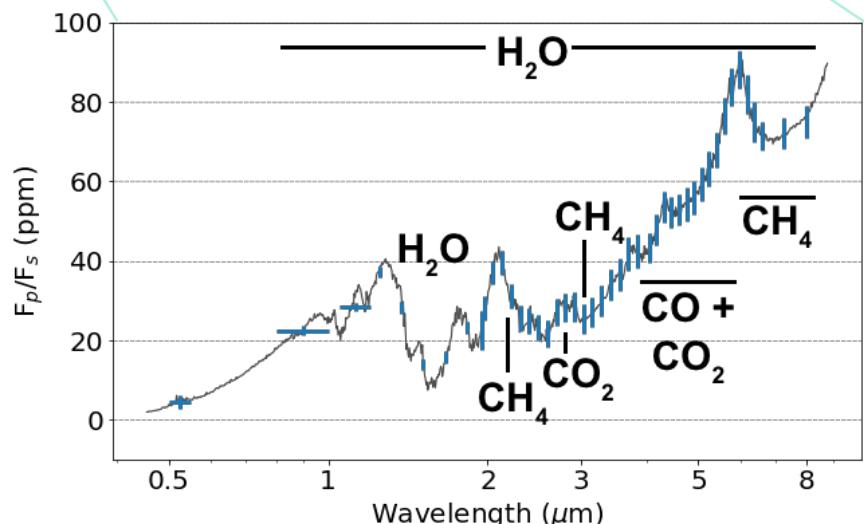
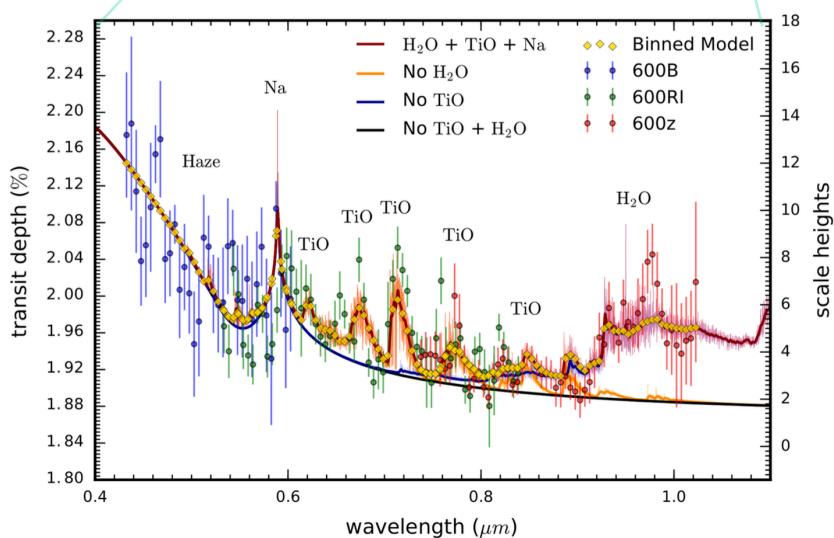
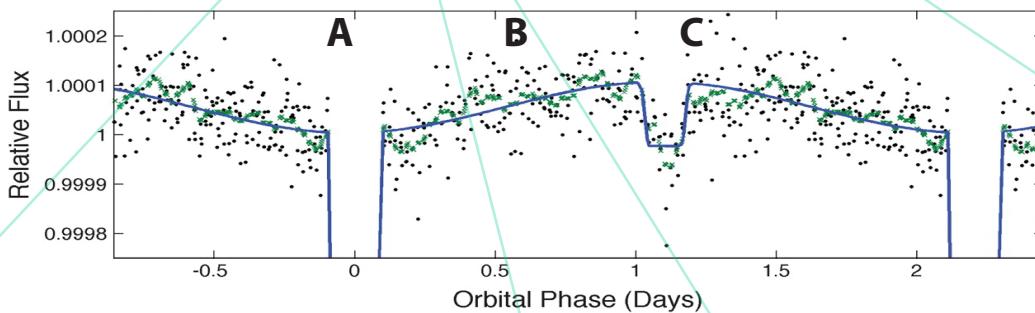
Channel Name	Wavelength ( $\mu\text{m}$ )	Spectral Resolution
VisPhot	0.5 – 0.55	Photometer
FGS-1	<b>0.6 – 0.8</b>	Photometer
FGS-2	<b>0.8 – 1.1</b>	Photometer
NIRSpec	<b>1.1 – 1.95</b>	$R \sim 20$
AIRS-Ch0	1.95 – 3.9	$R = 102 – 180$
AIRS-Ch1	3.9 – 7.8	$R = 30 – 64$

# The exoplanet mission timeline



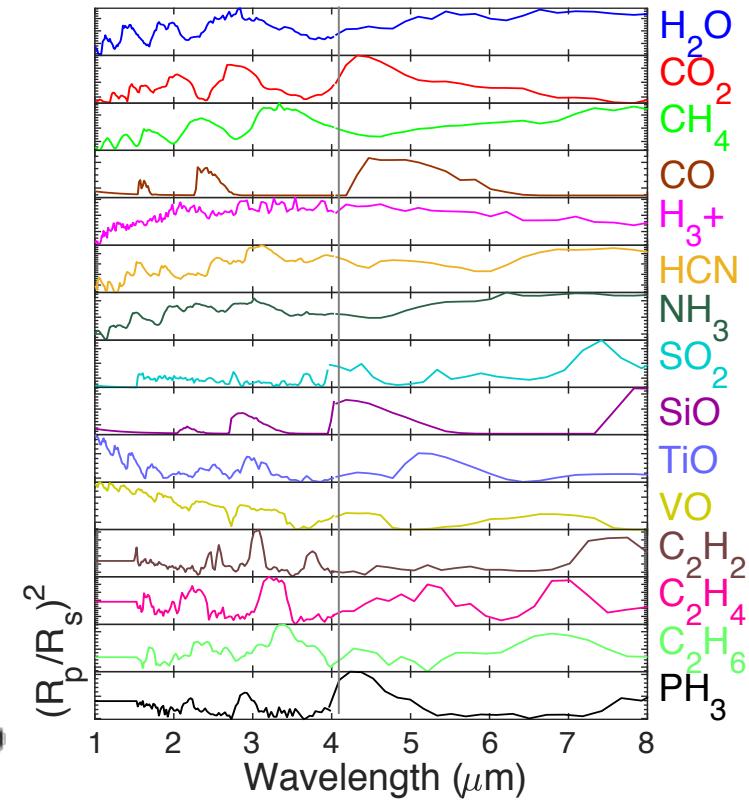
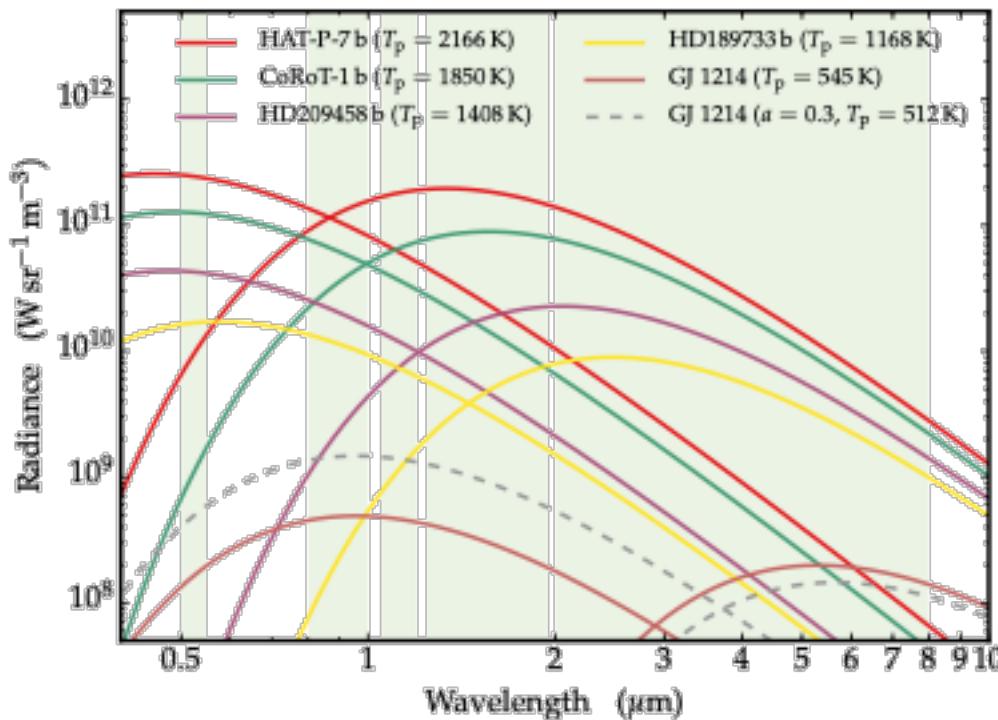


$10^{-4} - 10^{-5}$  relative precision



# A chemical survey of a large population

SCIENCE REQUIREMENTS: **EXOPLANET RADIATION, MOLECULAR & CLOUD SIGNATURES, STELLAR ACTIVITY**

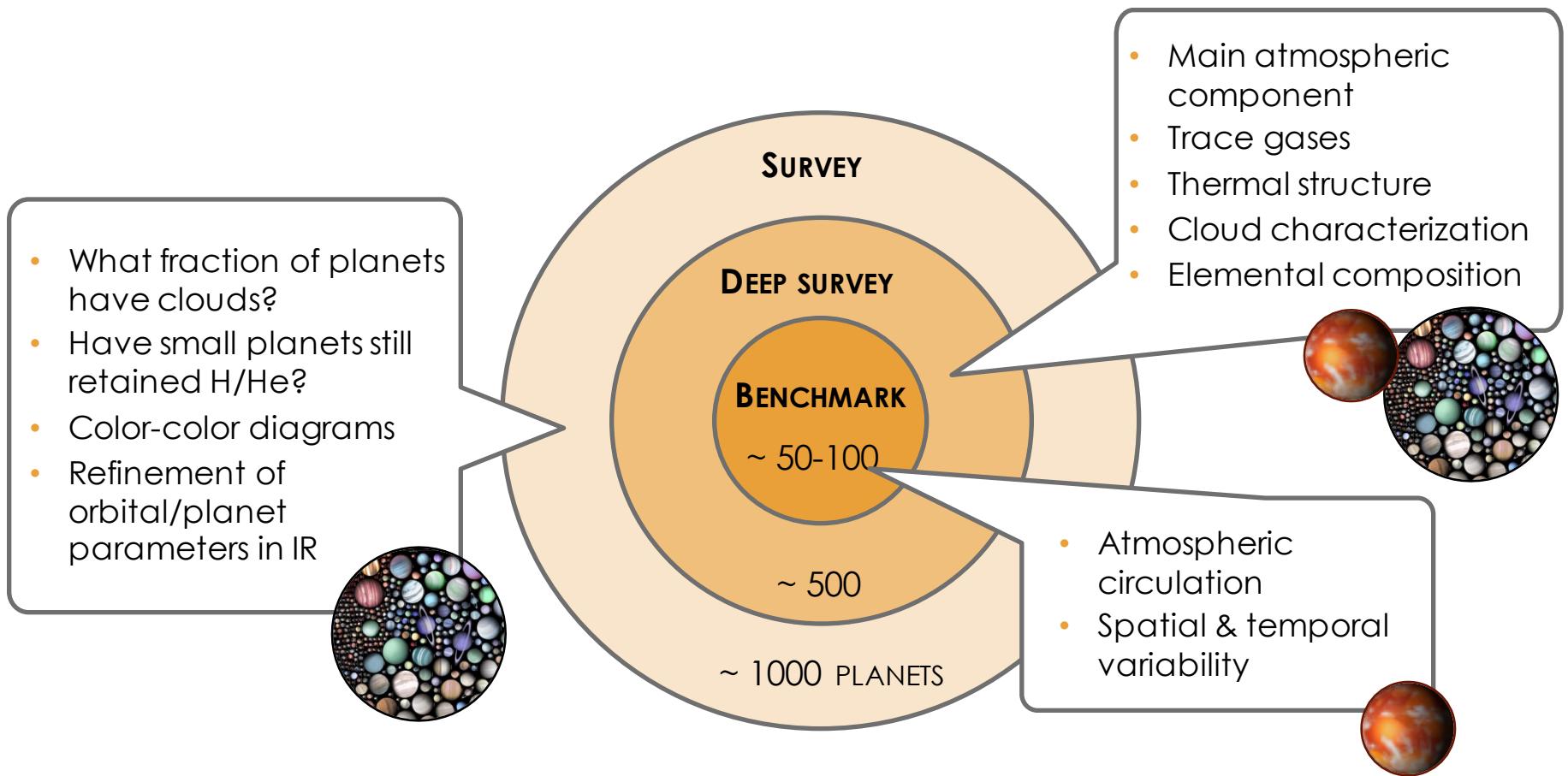


Simultaneous observations in the VIS and IR are needed



# ARIEL 3-tier approach

## INDIVIDUAL PLANETS & POPULATION ANALYSIS



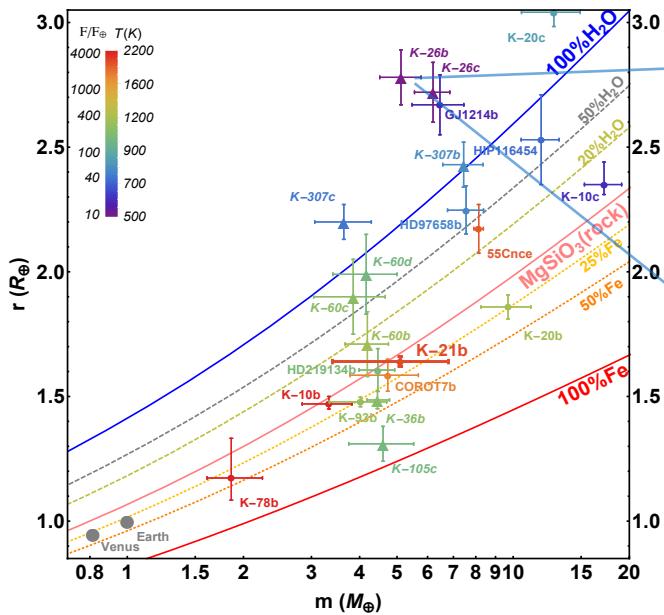


# Chemical diversity

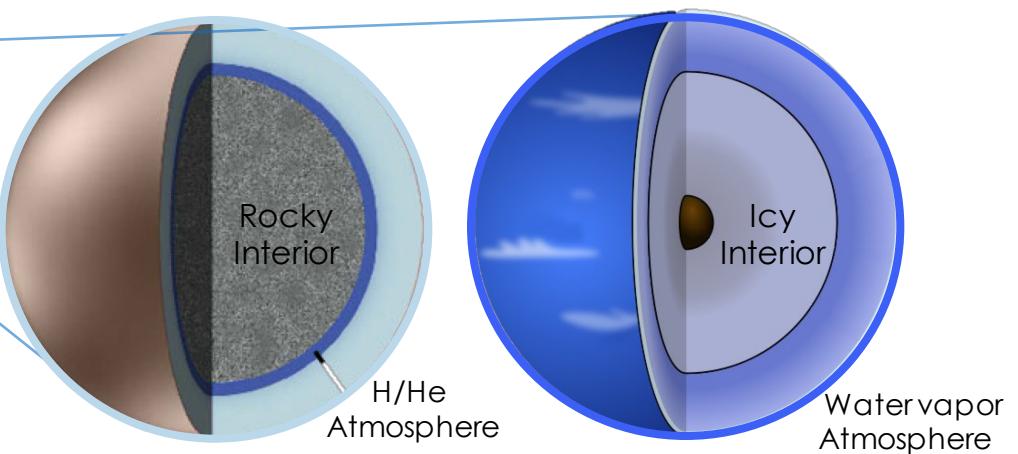


ARIEL WILL CLARIFY CORRELATION WITH THE DENSITY

## Density observations



Atmospheric composition through  
ARIEL will clarify the degeneracy

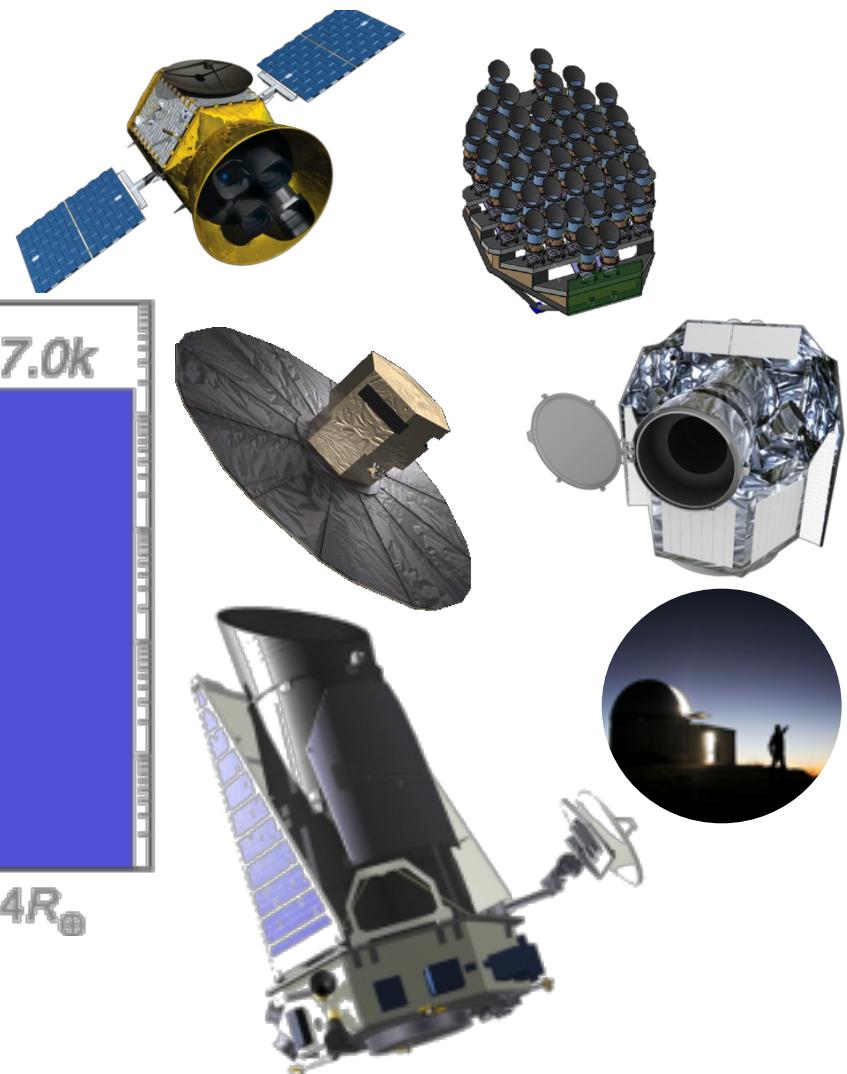
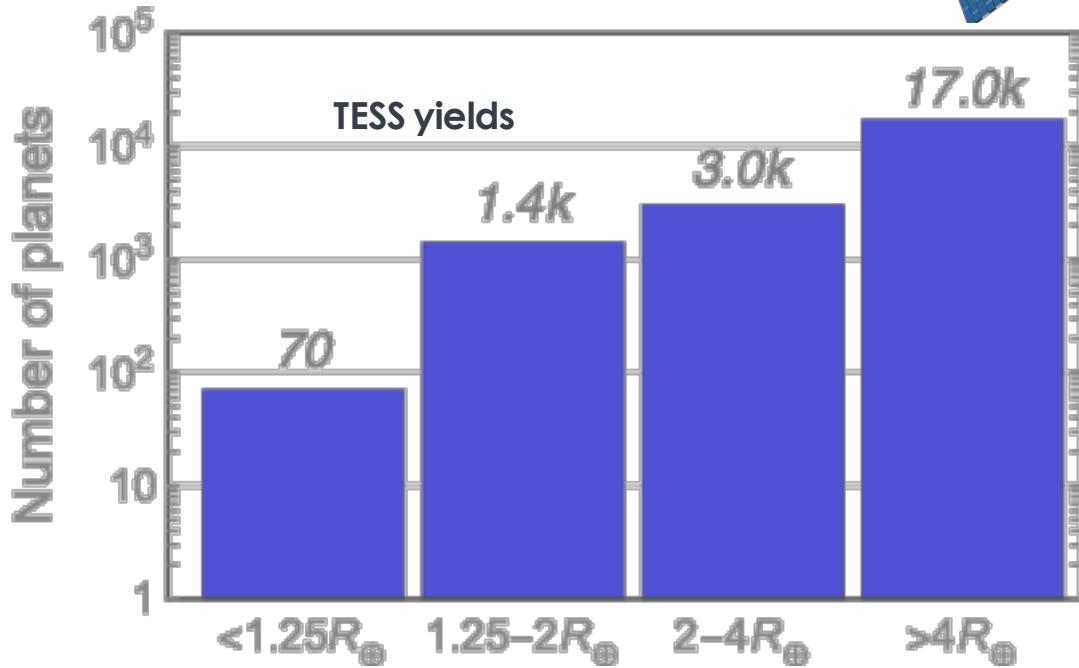


Same mean density – Different atmospheric signatures

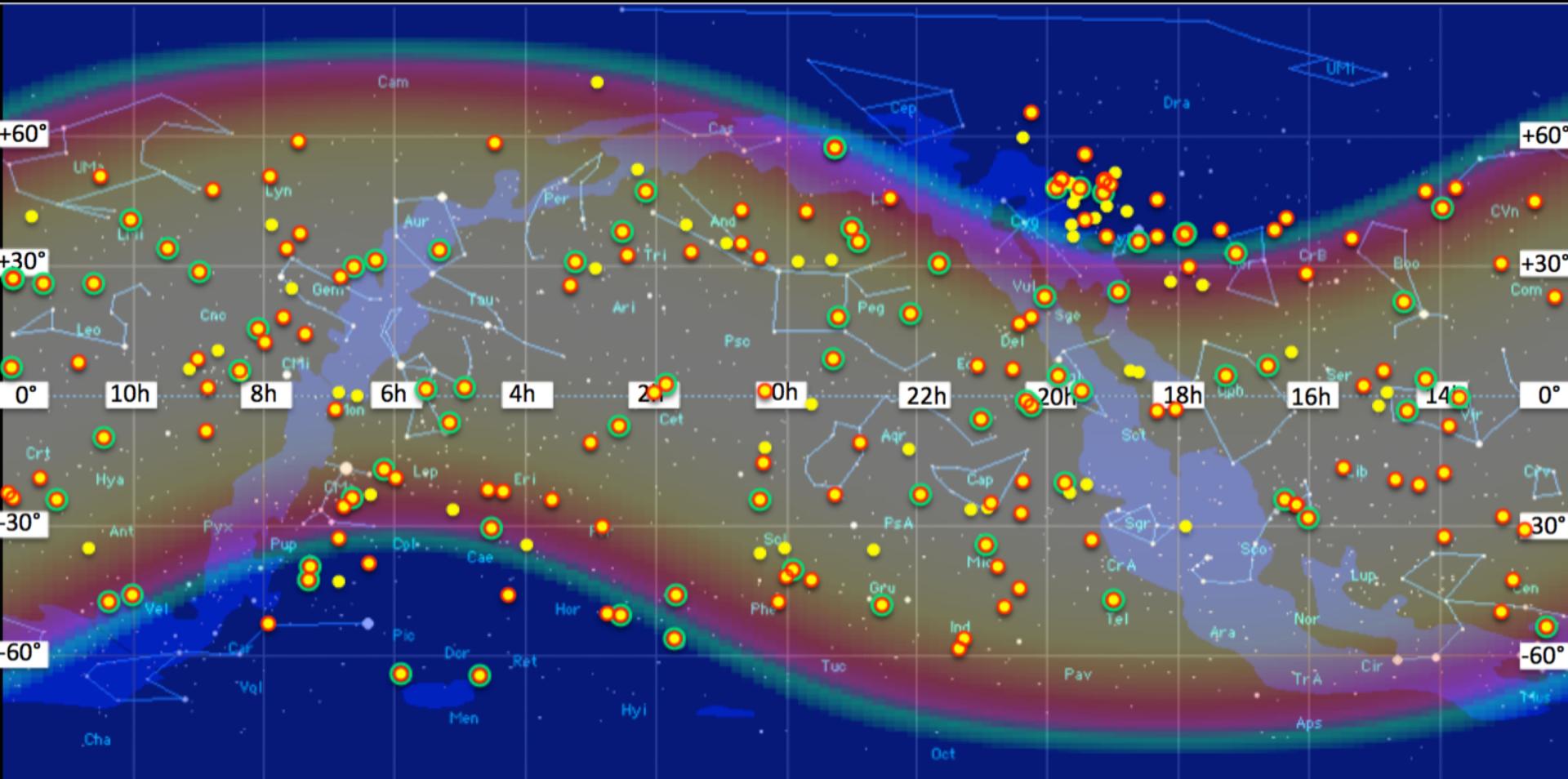
# *Large population of warm/hot planets*



TODAY AND IN THE NEXT DECADE



# ARIEL sky visibility



Instantaneous Sky  
Visibility



Survey  
Deep  
Benchmark

# *ARIEL science working groups*



## **Spectral retrieval:**

Michiel Min (SRON), Joanna Barstow (UCL), Ingo Waldmann (UCL)

## **Planet interior:**

Ravit Helled (Un. Zurich), Stephanie Werner (Un. Oslo)

## **Planet formation:**

Diego Turrini (INAF)

## **Disks:**

Mihkel Kama (Cambridge), Olja Panic (Leeds)

## **Atmospheric chemistry:**

Olivia Venot (LISA), Yamila Miguel (Un. Leiden), Franck Selsis (Un. Bordeaux)

## **Mass measurement with RV:**

Lars A. Buchhave (DTU), Olivier Demangeon (Un. Porto)

## **Stellar characterisation:**

Camilla Danielski (CEA)

## **Data challenges:**

Subi Sarkar (Un. Cardiff) and Theresa Leuftinger (Un. Vienna) (Stellar activity)

Nikos Nikolaou (UCL) (Data reduction), Michel Min (SRON) (Retrievals)

## **Stellar activity:**

Bart Vandenbussche (Un. Leuven), **Ignasi Ribas (IEEC-CSIC)**, Giusi Micela (INAF)

## **Data analysis:**

Ingo Waldmann (UCL), Angelos Tsiaras (UCL)

# ***ARIEL science working groups***



## **Upper atmosphere/star-planet interaction:**

Antonio Garcia Muñoz (TU Berlin), Manuel Guedel (U. Vienna), Luca Fossati (U. Graz)

## **High-cadence photometry:**

Davis Waltham (RH), Gyula Szabo (Un. Eötvös Loránd), Luca Borsato (Un. Padova)

## **Ephemerides:**

Vincent Coudé du Foresto (Obs. Paris)

## **Phase-curves:**

Benjamin Charnay (Obs. Paris), João Mendonça (DTU)

## **Albedo and reflected light (+ synergy with other instruments):**

Olivier Demangeon (Un. Porto)

## **Synergy with ELT & ground-based obs.:**

**Enric Palle (IAC)**

## **Synergy with JWST:**

Pierre-Olivier Lagage (CEA)

## **Synergy with Solar System science:**

Gabriella Gilli (Obs. Lisboa), Pedro Machado (Obs. Lisboa)

## **Synergy with PLATO and CHEOPS:**

Isabella Pagano (INAF), Giampaolo Piotto (Un. Padova)

# ***ARIEL science working groups***



## **Scheduling:**

Juan Carlos Morales (ICEE-CSIC), Andrea Moneti (IAP)

## **Performance simulations:**

Enzo Pascale (La Sapienza/Cardiff), Subi Sarkar (Cardiff), Billy Edwards (UCL), Lorenzo Mugnai (La Sapienza)

## **Spectroscopic database:**

Clara Sousa-Silva (MIT), Svatopluk Civiš (Heyrovsky Institute)

## **Other science: brown dwarfs:**

Sarah Casewell (Un. Leicester)

## **Other science: young stellar objects:**

Csaba Kiss (Konkoly Obs.), Krisztián Vida (Konkoly Obs.)

## **Novel methods to handle high-volume, high-dimensional data:**

James Cho (QMUL)

## **Synergy with NGTS, HAT:**

Matt Burleigh (Leicester Un.), Gaspar Bakos (Princeton)

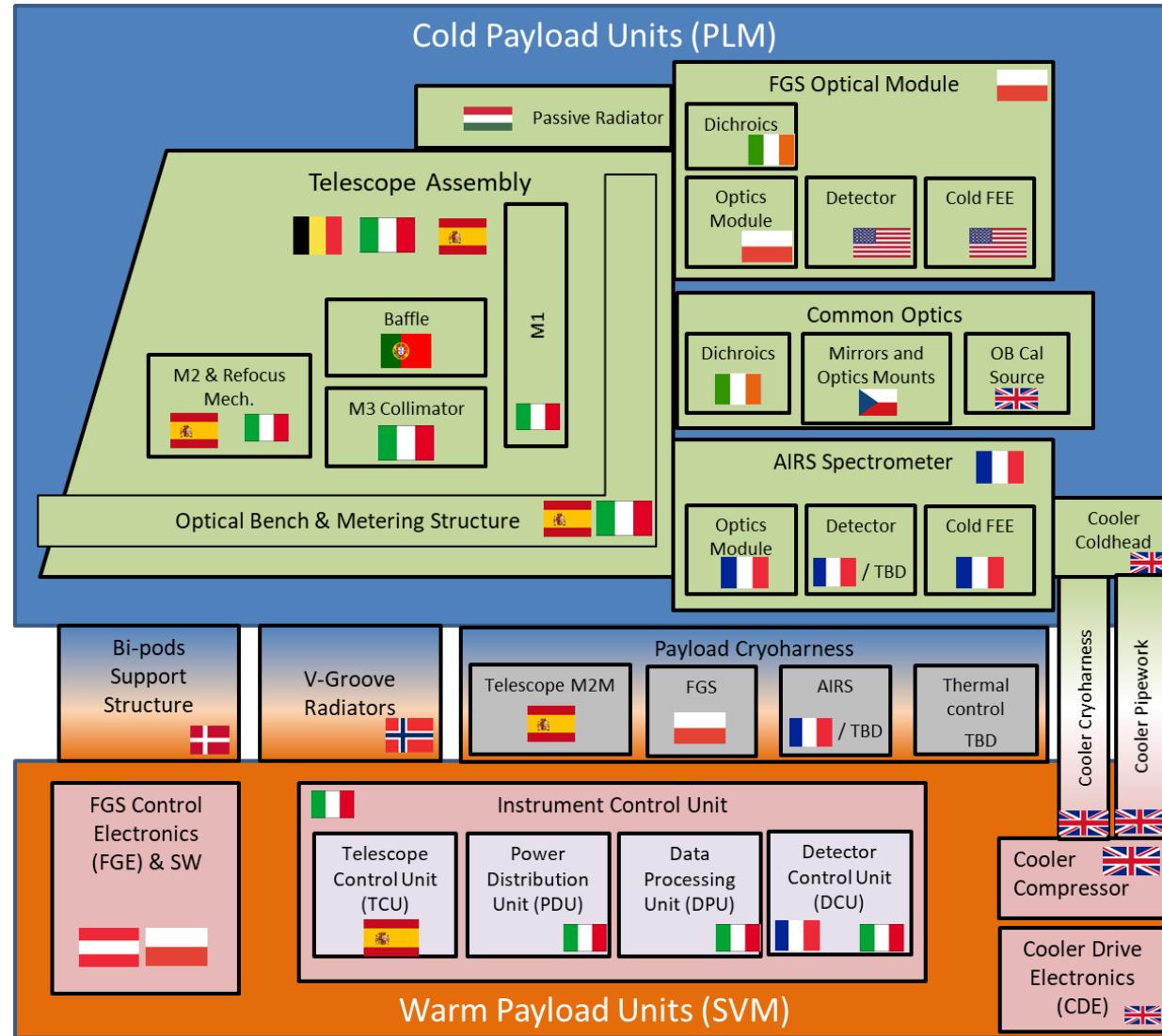
## **Synergy with TESS:**

TBA from US

## **Synergy with amateur astronomers:**

Anastasia Konkori (Royal Museums Greenwich), Marco Rocchetto (Konica Minolta)

# ARIEL payload



+ EGSE  
+ mission planning

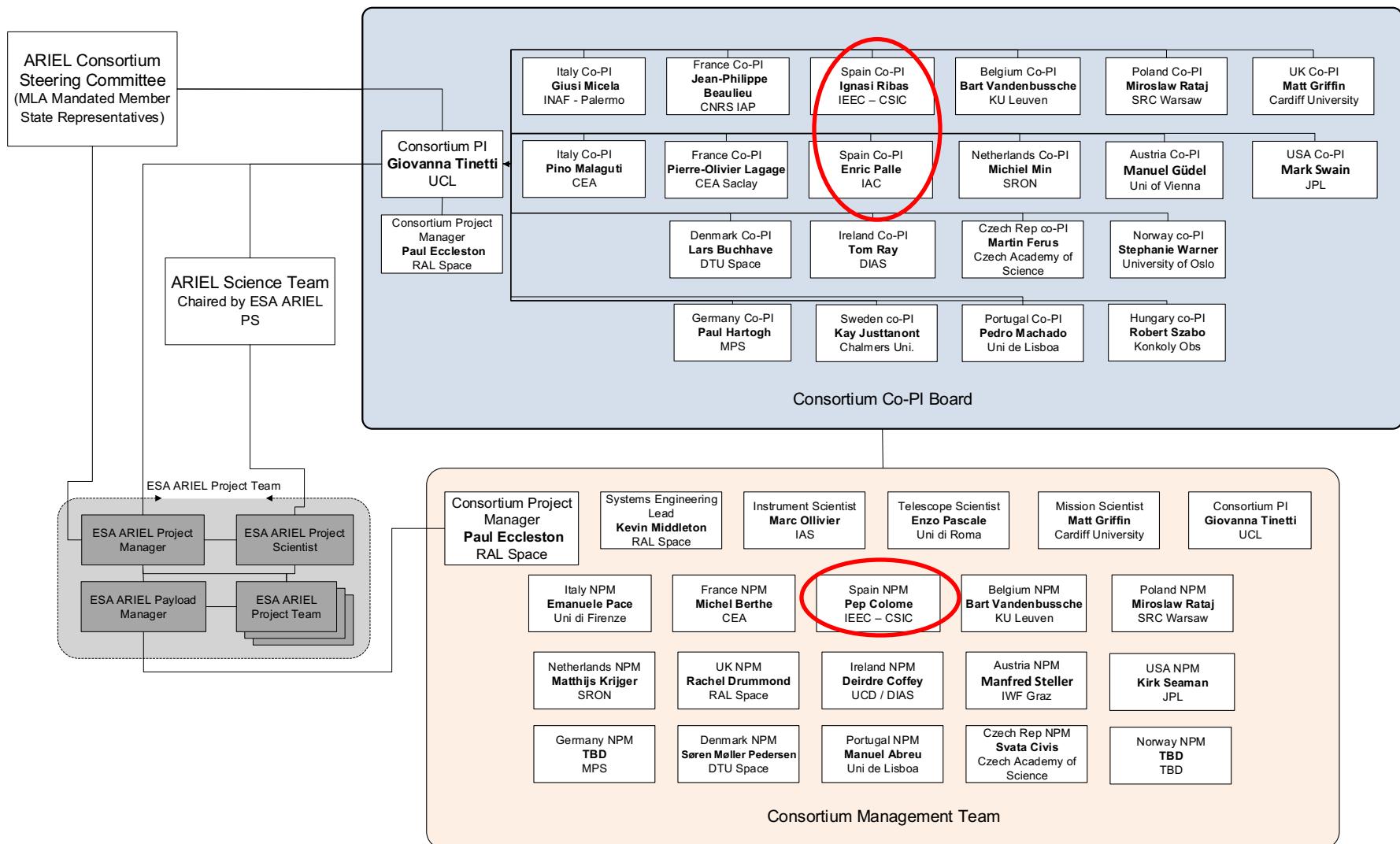
Telescope Assy  
ALV & Cryo-  
testing  
[Belgium flag]

Payload Overall ALV, Env. Test &  
Calibration  
[UK flag]  
PLM Level OGSE [UK, Portugal flags]  
PLM Level MGSE [Hungary flag]

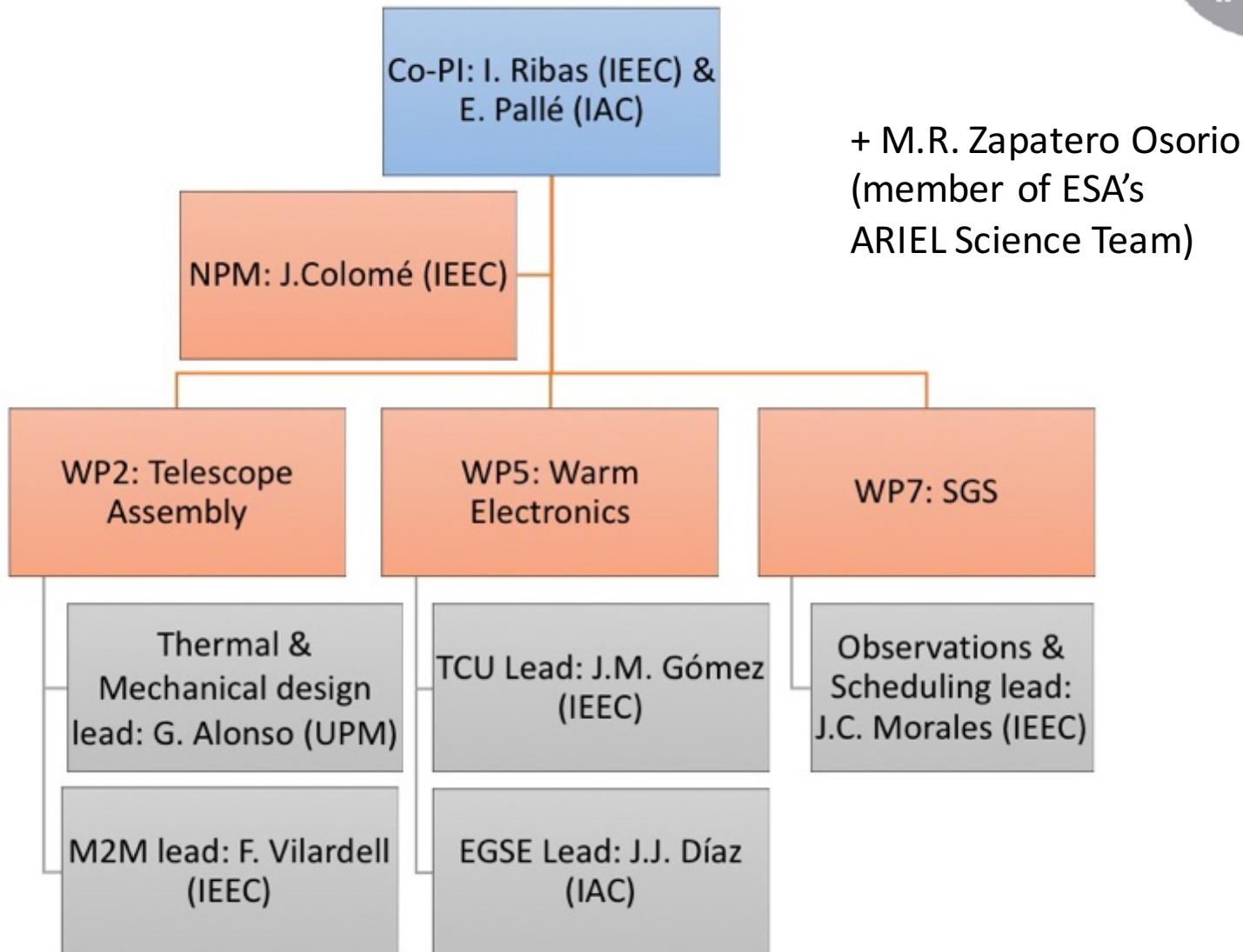
Systems Team  
Leadership  
[UK, France, Italy, Belgium, Poland flags]

Consortium  
Management  
& Coordination  
[UK flag]

# *Our role in the ARIEL consortium*



# *Spanish consortium organization*



# *Conclusions*



- ARIEL has been conceived to deliver the first chemical survey of ~1000 exoplanets, probing uniformly the gamut of planet and stellar parameters
- The Spanish participation in the ARIEL mission science consortium and payload development is very significant
- Timeline: Phase B1 (2019-2020), mission adoption (Nov 2020), phase B2 (2021-2022)
- Think about joining the science WGs!

